2^{ème} Ecole d'été GRGS "Géodésie spatiale, physique de la mesure et physique fondamentale".

30 août - 4 sept 2004 - Forcalquier (France)

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Gravity probe B experiment, **Superconducting gyro** & atomic gyros

Precision Clocks in Space and GPA H-maser (1976)

Gravity Probe A (1976)

Vessot et al, PRL 45, 2081 (1980)

• Comparison of two clocks at different gravity potential

• On ground and on board a rocket with parabolic trajectory (10 000 km max. altitude)

• Redshift of 4x10⁻¹⁰ measured with a 10⁻¹⁴ clock frequency stability

• 70 ppm confirmation of combined redshift and 2nd order Doppler

ACES/PHARAO (ISS : 2008 ? Or other S/C : ?) expected accuracy : 25 better



, DMPH

GRAVITY PROBE B Scientific Objectives



septembre 2004 4 1 30 août Ecole d'été Géodésie spatiale 2 è Toubou ф **2 GRAVITY PROBE**

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In orbit configuration



Circular Polar Orbit :

- Altitude : 640 km
 - 1-2 10⁻³
- Eccentricity :Inclination :

90.007 °

18 months operation (16 months present evaluation)

1 telescope

- 4 gyros (0.3 marcsec/year resolution) 1 GPS receiver Mass trim mechanism
- Mass trim mechanis
- 12 thrusters

- Star reference frame
- Ultra-sensitive gyros
- No disturbation
- Integration of the signal

CONTROLLED SPACE ENVIRONMENT with drag-free satellite orbit and cryogenic experiment : • Magnetic shielding

- Magnetic shielding
- Squid rotation detection
- Low thermal noise
- He thrust

Project Timeline : The basis

- **1893** Mach's Principal -The Science of Mechanics- acceleration relative to distant stars.
- **1887** Michelson & Morley Experiment : speed of light remains constant
- **1905** Einstein Special Relativity : propagation of matter and light at high speeds.

1915 Einstein General Relativity : gravitational forces in terms of space curvature caused by the presence of mass.

Fundamental principle : accelerated frames and in gravitation fields frames are equivalent. General Relativity predicts : clocks evolution in gravitational fields (or accelerated frames), gravitational redshift, existence of gravitational lensing, gravitational waves, gravitomagnetism, Lense-Thirring effect, and relativistic precession of orbiting bodies.

1924 J. Lense and H. Thirring

calculated effect : a rotating object will slowly drag space and time around with it! A moon orbiting a rotating planet undergoes a relativistic advance of its ascending node. Frame Drag.

1929 A. S. Eddington : proposed an Earth based gyroscope or pendulum experiment of general relativity.

If the earth's rotation could be accurately measured by Foucault's pendulum or by gyrostatic experiments, the result would differ from the rotation relative to the fixed stars by this amount of 19 milliarcsecond/year precession.

Project Timeline : The Fondation

- **1961** First formal NASA contact : Fairbank writes Dr Abe Siberstein describing an instrument that would measure the geodetic precession to a few percent.
- **1962** Francis Everitt joins William Fairbank and Leonard Shiff at Stanford on the Gravity ProbeB.
- **1965** 1st fused quartz telescope built.
- **1971** NASA begins examining feasibility of a flight experiment. Ball Aerospace completed a Mission Definition Study.
- **1973** Dan Debra's successful flight of a drag-free satellite (the Transit navigation satellite).
- **1976** Gravity Probe A launch. 1 hour 55 minute flight of a MASER atomic clock demonstrating time change as weaker levels of gravity : test of redshift to an accuracy of 2.10⁻⁴.
- **1977** End of longest single continuous research NASA grant ever awarded (63-77).

1980-82 Phase A at MSFC leading to larger dewar and satellite.

Project Timeline : The mission happens

- **1983** Stanford restructured program : science instrument within the dewar to be integrated and launched in 1991 on the shuttle : STORE (Shuttle Test of the Relativity Experiment)
- **1985** Gyro production throws out Beryllium, Hollowed Beryllium, Hollow Quartz spheres and focuses on Quartz rotors...
- **1986** Challenger explodes.
- **1989** Stanford's first prolonged levitation of a quartz sphere.
- **1992** First Flight Hardware within the Science Mission starts to be built : Dewar...
- **1995** NASA cancels Shuttle Test and directs Stanford to go directly to flight.
- **2001** Integrate Payload with Spacecraft.

April 20th 2004Gravity Probe B successfull launch
out of Vandenburg Air Force Base at 9:55am.

• The Satellite



VEHICLE





Payload general configuration

- From cryogenic (He liq. 1.8 K) to room temperature
- Alignment : Telescope, Gyros, S/C spin axis
- Drag free satellite : 10⁻⁹ g
- S/C mass centring
- Satellite rotation : ~ 10⁻² Hz (period : 1 to 3 mn)



4 gyros for redundancy and performance improvement drift rate : 0.25 marsec/year leads to accuracy $\gamma \approx 2.10^{-5}$ $\alpha \approx 3 \ 10^{-3}$

Performance Payload with In the Dewar

PAYLOAD and **DEWAR**

- 2 441 liters of supercooled helium at 1.8 Kelvin (-271.4 C)
- 2.74 m tall / 2.64 m diameter
- Porous plug at the top : as the internal liquid helium heats up, it evaporates and the gas is vented out taking heat with it.
- Payload Power Usage: 313 Watts





- 2^{éme} Ecole d'été Géodésie spatiale - 30 août - 4 septembre 2004 -

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The Quartz Block





QUARTZ BLOCK

- Weight : 34 kg
- Length : 55 cm
- Diameter : 18.5 cm
- Block lapped and polished (14 months to hand-polish)
- Telescope mounting surface of the block had to be poliched to within 0.01 um
- polished to within 0.01 µm





The Gyroscopes



GYROSCOPE

- Ball (rotor) size- 3.81 centimeter diameter (1.5-inch)
- Homogeneous fused quartz : 2 10⁻⁶
- Sphericity : less than 40 atomic layers from perfect (1nm)
- Coating- Niobium (uniform layer 1,270 nanometers thick)
- Electrostaticaly suspended (25 µm gap).
- Spin Rate- Between 5,000 and 10,000 RPM (obtained once by He flow)
- Accuracy : 0.25 marcsec/year drift (0.5 10⁻¹⁶ rd/s)

• Friction





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The SQUID's rotation measurement

SQUID's

- Cryogenic magnetic field variation sensor.
- Superconducting loop with 2 Josephson junctions
- Sensitivity : 5x10⁻¹⁴ gauss (5x10⁻¹⁸ Tesla) 10⁻¹³ of the Earth's magnetic field.

Rotation Measurement :

London Effect

6.10^{-12°}/ hour (< 10⁻⁶ best nav. gyro performance)





London effect induces magnetic moment the variation of orientation is detected by SQUID





• The Telescope



CASSEGRAIN TELESCOPE

- Composition- Homogeneous fused quartz
- Length 35.56 centimeters (14 inches)
- Aperture 13.97 centimeter (5.5-inch)
- Focal length 3.81 meters (12.5 feet)
- Mirror diameter 14.2 centimeters (5.6 inches)
- Guide Star HR 8703 (IM Pegasi : Mag 5.6)
- Accuracy : 0 .1 milliarcsecond i.e. 5.10⁻¹⁰ rd



The star tracker & Sunshade







SUNSHADE

Inside sun shield : series of black, metal baffles to absorb incoming stray light before it can reach the telescope.

STAR TRACKER

Two star trackers : wide field and narrow field (star sensor). Star sensor : field of view ~ 1° (1.7 10^{-2} rd) resolution \sim 1 arcminute (3 10⁻⁴ rd) in GP-B telescope field of view,

-> Guide star's position to 1 milliarcsecond (5 10-9 rd).



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E The GMA

The Gas Management Assembly



GMA FACTS

- Helium gas (99.999% pure) used to spin up the gyroscope ball.
- Helium gas used for thrusters of the drag free control.
- Fine distribution and management of the evaporated He to be ejected from the dewar



The Thrusters

THRUSTER

- 12 pairs of thrusters on the vehicle.
- Use of the evaporated liquid helium from the dewar as a propellant linear thruster independent of the inlet pressure

Objective :

- Fine control of the satellite attitude and orbit
- Satellite rotates to modulate the SQUID output (reduction of noise)







GPB Mission Present Status (Cospar july 04)

Satellite in nominal orbit and nominal operation

- Drag free and attitude control being optimised : telescope pointing not yet stabilized along reference star
- 2 gyros rotates at nominal frequency
- He Dewar : 14 months mission evaluated
- Calibration phase running : no scientific results before 6 months

Existing gyroscope technology for craft attitude motion application

ESGN (submarine navigation)

Draper LN-TGG gyro

Litton/Northrop Hemispherical Resonator



LN-TGG; 1 nrad 0.1-100 Hz Source: SPIE 4632-15



Fibersense IFOG

Atom interferometer force sensors

The quantum mechanical wave-like properties of atoms can be used to sense inertial forces.

Gravity/Accelerations

As atom climbs gravitational potential, velocity decreases and wavelength increases

Rotations

Rotations induce path length differences by shifting the positions of beam splitting optics





(Light-pulse) atom interferometry



Refroidissement et positionnement des atomes :

piège magnéto-optique (avec gradient de champs magnétiques)

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Interférométrie à onde de matière



Capteurs inertiels à ondes de matière

Interférométrie atomique : (ici le plus simple : MACH - ZENDER)



12 омрн





Capteurs inertiels à ondes de matière

Interférométrie atomique

Historique

Rési

- 1923 : Louis de Broglie
- 1927 : interférence sur e⁻ (Davisson- Germer)
- 1974 : expérience COW (neutron lent)
- 1980 : écho de photon (Mossberg)
- 1991 : Sodium (Kasevich / Chu)
- 2001 : Formalisme ABCD abouti (Bordé)

ltats	- Gyromètre (à jet thermique) : Kasevich 1991	1: 6	.10 ⁻¹⁰	rad	s/\sqrt{Hz}
	- Gravimètre (à atomes froids) : Chu 1999 :	5.10 3	$5^{-9} g$.10 ⁻⁹	ou g sui	2.10 ⁻⁸ g r 60s
	- Gradiomètre : Kasevich 2000 :	4.	.10 ⁻⁹	s^{-2}/r	\sqrt{Hz}

En France, développements au SYRTE, IOTA, LPL, ONERA

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Al inertial sensors: performance summary

	Demonstrated ground	Anticipated ground	Projected space
Gyroscope ARW Bias stability Scale factor	2x10 ⁻⁶ deg/hr ^{1/2} 6x10 ⁻⁵ deg/hr 5 ppm	<1x10 ⁻⁶ deg/hr ^{1/2} <10 ⁻⁵ deg/hr <1 ppm	<10 ⁻⁸ deg/hr ^{1/2} <10 ⁻⁷ deg/hr <1 ppm
Accelerometer Sensitivity Bias stability Scale factor	10 ⁻⁹ g/Hz ^{1/2} <10 ⁻¹⁰ g <10 ⁻¹⁰	<10 ⁻¹⁰ g/Hz ^{1/2} <10 ⁻¹⁰ g <10 ⁻¹⁰	<10 ⁻¹³ g/Hz ^{1/2} <10 ⁻¹⁶ g ? <10 ⁻¹²

Kasevitvh 2003

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Accelerometer Technology

Ground and Space Applications



Capteurs inertiels à ondes de matière



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Gravimètre



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Stanford/Yale laboratory gravity gradiometer

1.4 m



Distinguish gravity induced accelerations from those due to platform motion with differential acceleration measurements.



Demonstrated diffential acceleration sensitivity:

4x10⁻⁹ g/Hz^{1/2}

(2.8x10⁻⁹ g/Hz^{1/2} per accelerometer)

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Gravimètre

Calcul général et exact du déphasage ...

Formules disponibles : - pour les gradients - général en T et T ' - ...

which can be written to first-order in γ , with T=T':

$$\delta \varphi = kgT^2 + k\gamma T^2 \left[\frac{7}{12}gT^2 - \left(v_0 + \frac{\hbar k}{2M} \right)T - z_0 \right]$$

Reference: Ch. J. B., Theoretical tools for atom optics and interferometry, C.R. Acad. Sci. Paris, 2, Série IV, p. 509-530, 2001

ју смрн



Systèmes de Référence Temps-Espace

SYRTE



Measurement of G



Systematic	$\frac{\delta G}{G}$
Initial Atom Velocity	1.88×10^{-3}
Initial Atom Position	$1.85 imes 10^{-3}$
Pb Magnetic Field Gradients	$1.00 imes 10^{-3}$
Rotations	$0.98 imes 10^{-3}$
Source Positioning	$0.82 imes 10^{-3}$
Source Mass Density	$0.36 imes 10^{-3}$
Source Mass Dimensions	$0.34 imes 10^{-3}$
Gravimeter Separation	$0.19 imes 10^{-3}$
Source Mass Density inhomogeneity	$0.16 imes 10^{-3}$
TOTAL	3.15×10^{-3}

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Present sensitivity/accuracy: $dG = 3 \times 10^{-3} G$

Measurement consistent with accepted value

Д омрн

IMU

Full IMU/compact sensor arrangement:

- 10 μdeg/hr^{1/2} ARW (2.5 10⁻⁹ rd/s^{1/2})
- 10⁻⁸ g/Hz^{1/2} accel noise
- 100 Hz bandwidth
- Expected excellent bias stabilities for accel/gyro
- Laser/control electronics similar to gravity gradient sensor
- Robust





3 axes rotation + 3 acceleration



Accelerometry





THE END

Gravimètre

Perfo - calcul simpliste ... (exemple du cas A. Peters/ S. Chu)

	Lbdo =	957		_	k off -	14749261 2	no. 1						
		432			Ken –	2 24 5 25 1,3	111-1						_
	masse atome -	133	uma —		masse –	2,21E-25	кg		ag/g –	2,00E-08	UC al		
			atomes	~	Vrecui =	7,05E-03	m/s			20,0	MGai		
		1,5	μκ —		& thermique =	0,0096829	m/s						
	I lancement =	0,13	s		ox en fin =	11,8131342	mm	_					
	Tentre Raman =	0,16	s					_	a c				
	T détection =	0,16	S						0,5		1.		
	g =	9,8	m/s/s —		dphi(g)=	3700294,67			0.4				
	v lancement =	3	m/s		hmax =	0,45918367	m	_	0,4				
	detection							-	03				
	monocoup a	85						Ē	0,0				
	la frange sur							7	0.2				
	uma	1,66E-27	kg						0,1	_	_	\	
	k	1,38E-23	J/K						1				
	h	6,63E-34	J.s						o 🖌 ——			,	
	pi	3,14E+00							0	0,2	0	,4 (0,6
										te	emps (s)		
	temps	Pulses	√z	z	z-	z+							
Lancement	0		3	0	0	0							
	0,026		2,7452	0,07	0,074435845	0,07493936							
	0,052		2,4904	0,14	0,142246889	0,14325391							
	0,078		2,2356	0,2	0,203433134	0,20494367							
	0,104		1 ,9808	26, 0	0,257994579	26000862, 0							
	0,13	0	1,726	0,31	0,305931223	0,30844878							
Pulse π/2	0,13	0,505102	1,726	0,31	0,305931223	0,30844878							
	0,162		1,4124	0,36	0,355835771	0,35897303							
	0,194		1 ,0988	0,4	0,395705118	0,39946208							
	0,226		0,7852	0,43	0,425539265	0,42991593							
	0,258		0,4716	0,45	0,445338213	0,45033459							
	0,29	0	0,158	0,46	0,45510196	0,46071804							
pulse π	0,29	0,505102	0,158	0,46	0,45510196	0,46071804							
	0,322		-0,1556	0,46	0,454830507	0,46106629							
	0,354		-0,4692	0,45	0,444523854	0,45137935							
	0,386		-0,7828	0,43	0,424182002	0,4316572							
	0,418		-1,0964	0,4	0,393804949	0,40189985							
	0,45	0	-1,41	0,36	0,353392696	0,3621073							
pulse π/2	0,45	0,505102	-1,41	0,36	0,353392696	0,3621073							
	0,482		-1,7236	0,31	0,302945244	0,31227956							
	0,514		-2,0372	0,25	0,242462591	0,25241661							
	0,546		-2,3508	0,18	0,171944738	0,18251846							
	0,578		-2,6644	0,1	0,091391686	0,10258511							
	0,61		-2,978	0,01	0,000803433	0,01261657							
Détection	0.61		-2 978	0.01	0.000803433	0.01261657							

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Gravimètre

Points importants non abordés :

- déplacements lumineux (effet Stark AC/DC)
- asservissement du miroir Raman
- sources de bruit : bruit de phase haute fréquence, vibration, rotation, intensité laser, …

Noise source	S/N	σ ₂/μGal	
Atom shot noise (ð x 10⁶ atoms)	1700	0.16	
Detection noise	300	0.9	
Loran-C frequency stability	95	3.0	
Raman laser intensity noise	75	3.5	
Residual vibrations and rotations	55	5	
Background fluctuations	40	7	
High-frequency phase noise	25	11	
Overall known noise sources	18	15	
Observed noise	14	19	



ure 20. Effect of vibrations. (a) Relative sensitivity to different vibrational frequencies for an interferometer public (i) of a first of the sensitivity of the second relation of



- double interféromètre : même déphasage ..

12 омрн

Long distance acceleration measurement : « Pionner effect »



FIG. 1: Unmovelled acceleration as a function of distance from the sun, by Anderson *et al.* [3].



FIG. 8: CHASMP two-way Doppler residuals (observed Doppler velocity minus model Doppler velocity) for Pioneer 10 vs. time. 1 Hz is equal to 65 mm/s range change per second. The model is fully-relativistic. The solar system's gravitational field is represented by the Sun and its planetary systems [49].

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Gravity gradiometer: measurement of G



Pb mass translated vertically along gradient measurement axis.



Typical data:

~1x10⁻⁸ g change in acceleration due to gravitational forces for different Pb positions

<u>↓</u>2 дмрн

Prototype field ready sensor Sensor head Sensor optomechanics Laser system

Д2 омрн

Gyroscope

High sensitivity 2 axis



Gyroscope (~30 cm length, vertically oriented) ARW = 0.88 microdeg/hr1/2 Bandwidth 10 Hz





hbar = 1.054571596e-34; m = 2.2085e-025; k1 = 6*7.3833e+006; k2 = 6*7.3833e+006; k3 = 6*7.3833e+006; R = 6e6;	Propagation vector
g3 = -9.8; w2 = 7e-5; w3=7e-5; T_= 0.03;	Rotation rates (Earth rate for ref)
V3 = 5; Txx = 1.5e-6; Tyy = 1.5e-6; Tzz = -3.0e-6; Txy = 1.5e-6; Txz = 1.5e-6;	Gravity gradient
Tyz = $1.5e-6$; xg1 = .0001; xg2 = 1;	Position errors
xg3 = .0001; dv2 = .0001; dv3 = .0001;	velocity errors
dv1 = .0001; dk1 = 1e-6*k1; dk2 = 1e-6*k2; dk3 = 1e-6*k3;	Alignment error
Term -2*TA2*w2*k1*v3 -2*TA3*w2*g3*k1 TA2*dk3*g3 TA3*TX2*v3*k1 -2*TA3*k1*w2*R*w3A2 -2*TA3*k1*w2*R*w3A2 -2*TA3*k1*w2*R*w3A2 TA2*w2*R*w3*dk2 TA2*w2*R*dk3 -2*TA2*w3*dv2*k1 7/12*TA4*TX2*g3*k1 1/2/m*TA3*TX2*hbar*k1A2 -7/12*TA4*TX2*w2*w3*R*k1 7/12*TA4*TX2*w2A2*R*k1 -3/2/m*TA3*w3A2*hbar*k1A2 TA3*TX2*dv3*k1 TA3*TX2*dv3*k1 TA3*TX2*dv3*k1 TA3*TX2*dv2*k1 17/6*TA4*TZ2*k1*w2*v3 -17/6*TA4*TZ2*k1*w2*v3 -17/6*TA4*TZ2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*v3 -7/2*k1*w2*k1*w2*v3 -7/2*k1*w2*k1*w2*k1*w2*v3 -7/2*k1*w2*k1*w2*k1*w2*v3 -7/2*k1*w2*k1*w2*k1*w2*k1*w2*k1*w2*k1*w2*k1*	$\begin{array}{c c c c c c c c c c c c c c c c c c c $















The Solar Arrays

so·lar cell

def: A semiconductor device that converts the energy of sunlight into electric energy. Also called a photovoltaic cell.



SOLAR ARRAYS FACTS

- Each panel is 3.5 meters long by 1.3 meters wide
- The release mechanism is made up of Nitinol rods, commonly called "memory metal". When the rods are heated, they change shape and release the panels.
- The 9,552 individual Gallium Arsinide solar cells have an effeciency of 18.5%
- The total power needed to run the entire satellite would barely power the average microwave.



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The Truss Structure 51

Truss

def: An engineered structure of short framing members, such as beams, chords, and diagonals, assembled into a rigid support structure.

TRUSS STRUCTURE FACTS

• The truss structure is made of aluminum alloy beams, heliarc welded at the joints. Mechanical joints were not stiff enough to maintain the satellite's critical geometry.

• The structure's "open" frame design exposes the dewar to space, improving heat radiation.

• Equipment is attached by self-integrated pallets. Individual subsystems can be removed without disassembling the entire space craft.





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